Interpretation and solution of the cosmological constant problem.

Stéphane Wojnow,

wojnow.stephane@gmail.com

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## Abstract:

The cosmological constant problem or vacuum catastrophe has long been a mystery of physics. We bring a solution and a simple interpretation.

It is sufficient to calculate the dark energy density parameter  $\Omega\Lambda$  at Planck time  $t_p$ , origin of our universe :

We first consider its expression in the Friedman equation:

$$\Omega_{\Lambda} = \frac{\Lambda c^2}{3H^2}$$

then

with  $t_H = 1/H$ ,

where  $t_H$  is Hubble time and H is Hubble constant :

$$\Omega \Lambda, t_p = 1/3 \Lambda c^2 t_p^2$$

The vacuum catastrophe =  $\Lambda / l_p$  <sup>-2</sup> =  $\Lambda l_p$  <sup>2</sup>

or when the vacuum catastrophe is express in energy density (J/m³) instead of m<sup>-2</sup>

$$\frac{\epsilon_{\Lambda}}{\epsilon_{\rm Pl}} = \frac{\frac{3\Lambda c^4}{8\pi G}}{F_{\rm Pl} \ l_{\rm Pl}^{-2}}$$

where  $F_p = c^4/G$  is the Planck force

as

$$l_p = c t_p$$

$$l_p^2 = c^2 t_p^2$$

The vacuum catastrophe =  $\Lambda c^2 t_p^2$ 

The vacuum catastrophe =  $3 \Omega \Lambda, t_p$ 

## Conclusion

The vacuum catastrophe would be the energy density parameter of cosmological constant at Planck time, i.e. at the origin of the universe, in the  $\Lambda CDM$  model with a factor of 3 (and with a divisor of 8 pi if we express the problem in terms of energy density,  $J/m^3$ ), and it would no longer be a problem

References:

S.E. Rugh and H. Zinkernagel, The Quantum Vacuum and the Cosmological Constant Problem , <a href="https://arxiv:hep-th/0012253">arXiv:hep-th/0012253</a>

For the value  $l_p^{-2} = 3.83 * 10^{69} \, \mathrm{m}^{-2}$  from the QFT Lucas Lombriser, université de Genève, communiqué de presse , https://www.unige.ch/communication/communiques/2019/cosmologie-une-solution-a-la-pire-prediction-en-physique/